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## Observation and Spin-Parity Determination of the X(1835) in $J/\psi \rightarrow \gamma K_{S}^{0}K_{S}^{0}\eta$

M. Ablikim et al. (BESIII Collaboration)

Phys. Rev. Lett. **115**, 091803 — Published 27 August 2015

DOI: 10.1103/PhysRevLett.115.091803

## Observation and Spin-Parity Determination of the X(1835) in $J/\psi \to \gamma K_S^0 K_S^0 \eta$

Observation and Spin-Parity Determination of the X (1835) in J/ψ → γK<sup>0</sup><sub>8</sub>K<sup>0</sup><sub>8</sub>η M. Ablikim¹, M. N. Achasox<sup>0,1</sup>, X. C. Al¹, O. Albayrak², M. Albrecht⁴, D. J. Ambrose⁴, A. Amorose<sup>38,A,68C</sup>, F. F. An¹, Q. An⁴ose, J. Z. Bai¹, R. Baldini herroli³0.4, Y. Ban³¹, D. W. Bennett¹, D. J. Ambrose⁴, A. Amorose<sup>38,A,68C</sup>, F. F. An¹, Q. An⁴ose, J. Z. Bai¹, R. Baldini herroli³0.4, Y. Ban³¹, D. W. Bennett¹, J. V. Bennett⁴, M. Bertani³0.4, D. Bettoni³1.4, J. M. Bian³5 F. Bianchi³6.46° E. Boger³3.4, I. Boyko³3 R. A. Briere⁵ R. L. Cafo®, X. Cal¹-q. O. Cakit²0.40°, D. Cakit²0.40°, G. F. Cao¹, S. A. Cetin³0.9, J. F. Chang¹-a, C. Chelov³0.4c, G. Chelov ³1.4c, G. Chelov °1.4c, R. Chen², B. Y. Chen², J. C. Chen³, H. Y. Chen², J. C. Chen³, H. L. Chen², J. C. Chen³, M. L. Chen², S. J. Chen², X. R. Chen², Y. B. Chen¹, E. B. Chen¹, H. S. Chen², J. C. Chen³, H. L. Dal¹°, J. P. Dal³°, A. Divorsi³¹, D. Deoly\*\* C. J. V. Dong¹, M. J. Dong³°, M. Destofanis\*0.4ecc? F. P. De Mon³\*4.46°, Y. Diug²¹, C. Dong³°, J. Dong³°, L. D. Dong³°, S. Dong³°, J. Dong³°, M. Destofanis\*0.4ecc? F. P. De Mon³\*4.46°, Y. Diug²¹, C. Dong³°, J. Dong³°, S. Dong³°, J. Eng²³, D. E. Florvanati².4d. M. P. Ersteh¹\*2d. C. D. Pa¹, Q. Goo. Y. Y. Foug², L. Fuosi³2.4cc, F. Feldbauer²², G. Eelic?³\*0. C. Q. Fong³\*0. J. Eng²³, C. Con³³. S. Foug³\*1. Can², C. Con³³. S. E. Florvanati².4d. M. Pitsch¹\*2d. C. D. Pa¹, Q. Goo. Y. Y. Goo. Y. Y. Goo? Y. Goo. J. Garai².2d. Al. Con²°, C. Eng³\*0. E. Florvanati².4d. M. Pitsch¹\*2d. C. D. Pa¹, Q. Goo. Y. Y. Goo. Y. P. Goo. Y. Con³\*0. Z. Goo. S. Goo. J. Garai².2d. K. L. Be¹, C. L. B. Guo? Y. Y. Goo. Y. Con³\*0. Z. Goo. S. Goo. J. Garai².2d. K. L. Be¹, C. L. Ho²\*0. C. Ho²\*0. H. Hong³\*0. Y. J. Han³\*0. Y. J. Haa¹\*0. Y. Goo. Y. Goo. Y. L. Haa¹\*1. Y. L. Haa¹\*1. Y. L. Haa¹\*1. Y. C. Goo. Y. Goo. Y. L. Haa²\*1. Y. L.

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We report an observation of the process  $J/\psi \to \gamma X(1835) \to \gamma K_S^0 K_S^0 \eta$  at low  $K_S^0 K_S^0$  mass with a statistical significance larger than  $12.9\sigma$  using a data sample of  $1.31 \times 10^9$   $J/\psi$  events collected with the BESIII detector. In this region of phase space the  $K_S^0 K_S^0$  system is dominantly produced through the  $f_0(980)$ . By performing a partial wave analysis, we determine the spin-parity of the X(1835) to be  $J^{PC}=0^{-+}$ . The mass and width of the observed X(1835) are  $1844\pm 9(\text{stat})^{+16}_{-25}(\text{syst})$  MeV/ $c^2$  and  $192^{+20}_{-17}(\text{stat})^{+62}_{-43}(\text{syst})$  MeV, respectively, which are consistent with the results obtained by BESIII in the channel  $J/\psi \to \gamma \pi^+ \pi^- \eta'$ .

PACS numbers: 13.20.Gd, 13.66.Bc, 14.40.Be

The non-Abelian property of quantum chromodynamics (QCD) permits the existence of bound states beyond conventional mesons and baryons, such as glueballs, hybrid states and multiquark states. The search for these unconventional states is one of the main interests in experimental particle physics. One of the most promising candidates, the X(1835) resonance, was first observed in its decay to  $\pi^+\pi^-\eta'$  in the process  $J/\psi \to \gamma\pi^+\pi^-\eta'$  by BESII [1]; this observation was subsequently confirmed by BESIII [2]. The discovery of the X(1835) has stimulated theoretical speculation concerning its nature. Possible interpretations include a  $p\bar{p}$  bound state [3], a second radial excitation of the  $\eta'$  [4], and a pseudo-scalar glueball [5]. In addition, an enhancement in the invariant  $p\bar{p}$  mass at threshold,  $X(p\bar{p})$ , was first observed by BESII in the decay  $J/\psi \to \gamma p\bar{p}$  [6], and was later also seen by BESIII [7] and CLEO [8]. In a partial-wave analysis of  $J/\psi \to \gamma p\bar{p}$ , BESIII determined the  $J^{PC}$  of the  $X(p\bar{p})$  to be  $0^{-+}$  [9]. The mass of the  $X(p\bar{p})$  is consistent with the X(1835) mass measured in  $J/\psi \to \gamma \pi^+ \pi^- \eta'$  [2], but the width of the  $X(p\bar{p})$  is significantly narrower.

To understand the nature of the X(1835), it is crucial to measure its  $J^{PC}$  and to search for new decay modes. Because of its similarity to  $J/\psi \rightarrow \gamma \pi^+ \pi^- \eta'$ ,  $J/\psi \rightarrow$  $\gamma K \bar{K} \eta$  is a favorable channel to search for  $X(1835) \rightarrow$  $K\bar{K}\eta$ . In contrast to  $J/\psi \to \gamma K^+K^-\eta$ , there is no background contamination for  $J/\psi \rightarrow \gamma K_S^0 K_S^0 \eta$  from  $J/\psi \to \pi^0 K_S^0 K_S^0 \eta$  and  $J/\psi \to K_S^0 K_S^0 \eta$ , which are forbidden by exchange symmetry and  $\overline{CP}$  conservation. Therefore, the channel  $J/\psi \to \gamma K_S^0 K_S^0 \eta$  provides a clean environment with minimal uncertainties due to background modeling. In this Letter, we report the first observation and spin-parity determination of the X(1835) in  $J/\psi \to \gamma K_S^0 K_S^0 \eta$ , where the  $K_S^0$  and  $\eta$  are reconstructed from their decays to  $\pi^+\pi^-$  and  $\gamma\gamma$ , respectively. The analysis is based on a sample of  $(1310.6 \pm 10.5) \times 10^6 J/\psi$ events [10, 11] collected with the BESIII detector [12].

The BESIII detector is a magnetic spectrometer operating at BEPCII, a double-ring  $e^+e^-$  collider with center of mass energies between 2.0 and 4.6 GeV. The cylindrical core of the BESIII detector consists of a helium-based main drift chamber (MDC), a plastic scintillator timeof-flight system (TOF), and a CsI(Tl) electromagnetic calorimeter (EMC) that are all enclosed in a superconducting solenoidal magnet providing a 1.0 T (0.9 T in 2012, for about  $1087 \times 10^6$  collected  $J/\psi$ ) magnetic field. The solenoid is supported by an octagonal flux-return yoke with resistive plate counter muon identifier modules interleaved with steel. The acceptance of charged particles and photons is 93% of the  $4\pi$  solid angle, and the charged-particle momentum resolution at 1 GeV/c is 0.5%. The EMC measures photon energies with a resolution of 2.5% (5%) at 1 GeV in the barrel (endcaps). A GEANT4-based [13] Monte Carlo (MC) simulation software package is used to optimize the event selection criteria, estimate backgrounds and determine the detection efficiency.

Charged tracks are reconstructed using hits in the

MDC. Because there are two  $K_S^0$  with displaced vertices, the point of closest approach of each charged track to the  $e^+e^-$  interaction point is required to be within  $\pm 30$  cm in the beam direction and within 40 cm in the plane perpendicular to the beam direction. The polar angle between the direction of a charged track and the beam direction must satisfy  $|\cos\theta| < 0.93$ . Photon candidates are selected from showers in the EMC with the energy deposited in the EMC barrel region ( $|\cos\theta| < 0.8$ ) and the EMC endcaps region ( $0.86 < |\cos\theta| < 0.92$ ) greater than 25 MeV and 50 MeV, respectively. EMC cluster timing requirements are used to suppress electronic noise and energy deposits unrelated to the event.

Candidate  $J/\psi \rightarrow \gamma K_S^0 K_S^0 \eta$  events are required to have four charged tracks with zero net charge and at least three photon candidates. All charged tracks are reconstructed under the pion hypothesis. To reconstruct  $K_S^0$ candidates, the tracks of each  $\pi^+\pi^-$  pair are fitted to a common vertex.  $K_S^0$  candidates are required to satisfy  $|M_{\pi^+\pi^-} - m_{K_s^0}| < \tilde{0}.009 \text{ GeV}/c^2 \text{ and } L/\sigma_L > 2$ , where L and  $\sigma_L$  are the distance between the common vertex of the  $\pi^+\pi^-$  pair and the primary vertex, and its error, respectively. The  $\gamma\gamma\gamma K_S^0K_S^0$  candidates are subject to a kinematic fit with four constraints (4C), ensuring energy and momentum conservation. Only candidates where the fit yields a  $\chi_{4C}^2$  value less than 40 are retained for further analysis. For events with more than three photon candidates, multiple  $J/\psi \to \gamma K_S^0 K_S^0 \eta$  candidates are possible. Only the combination yielding the smallest  $\chi_{4C}^2$  is retained for further analysis. Candidate  $J/\psi \to \gamma K_S^0 K_S^0 \eta$ events are required to have exactly one pair of photons within the  $\eta$  mass window (0.51  $< M_{\gamma\gamma} < 0.57 \text{ GeV}/c^2$ ). Simulation studies show this criterion significantly reduces the miscombination of photons from 3.20% to 0.16%. The miscombination of pions is also studied and found to be negligible. To further suppress background events containing a  $\pi^0$ , events with any photon pair within a  $\pi^0$  mass window (0.10  $< M_{\gamma\gamma} < 0.16 \text{ GeV}/c^2$ ) are rejected. The decay  $J/\psi \to \phi K_S^0 K_S^0$  with  $\phi \to \gamma \eta$ leads to the same final state as the investigated reaction  $J/\psi \to \gamma K_S^0 K_S^0 \eta$ . Therefore, events in the mass region  $|M_{\gamma\eta} - m_{\phi}| < 0.04 \text{ GeV}/c^2$  are rejected.

After applying the selection criteria discussed above, the invariant mass spectrum of  $K_S^0 K_S^0 \eta$  shown in Fig. 1 (a) is obtained. Besides a distinct  $\eta_c$  signal, a clear structure around 1.85 GeV/ $c^2$  is observed. The  $K_S^0 K_S^0$ mass spectrum, shown in Fig. 1 (b), reveals a strong enhancement near the  $K_S^0 K_S^0$  mass threshold, which is interpreted as the  $f_0(980)$  by considering spin-parity and isospin conservation. The scatter plot of the invariant mass of  $K_S^0 K_S^0$  versus that of  $K_S^0 K_S^0 \eta$  is shown in Fig. 1 (c). A clear accumulation of events is seen around the intersection of the  $f_0(980)$  and the structure around 1.85  $\text{GeV}/c^2$ . This indicates that the structure around 1.85  $\text{GeV}/c^2$  is strongly correlated with  $f_0(980).$  By requiring  $M_{K_s^0K_s^0}<1.1~{\rm GeV}/c^2,$  the structure around 1.85  $\text{GeV}/c^2$  becomes much more prominent in the  $K_S^0 K_S^0 \eta$  mass spectrum (Fig. 1 (d)). In addition, there is an excess of events around 1.6  $\text{GeV}/c^2$ .

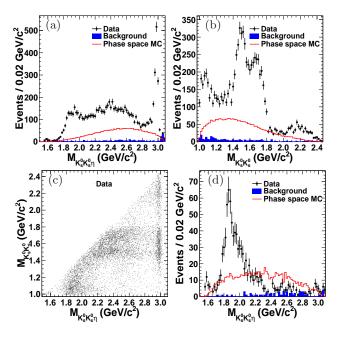


FIG. 1. Invariant mass distributions for selected events: Invariant mass spectra of (a)  $K_S^0K_S^0\eta$  and (b)  $K_S^0K_S^0$ ; (c) scatter plot of  $M_{K_S^0K_S^0}$  versus  $M_{K_S^0K_S^0\eta}$ ; (d)  $K_S^0K_S^0\eta$  invariant mass spectrum for events with the requirement  $M_{K_S^0K_S^0} < 1.1 \text{ GeV}/c^2$ . Dots with error bars are data; the shaded histograms are the non- $\eta$  backgrounds estimated by the  $\eta$  sideband; the solid histograms are phase space MC events of  $J/\psi \to \gamma K_S^0K_S^0\eta$  with arbitrary normalization.

Potential background processes are studied using a simulated sample of  $1.2 \times 10^9~J/\psi$  decays, in which the decays with measured branching fractions are generated by EVTGEN [14] and the remaining  $J/\psi$  decays are generated according to the LUNDCHARM [15] model. Simulated events are subject to the same selection procedure applied to data. No significant peaking background sources have been identified in the invariant mass spectrum of  $K_S^0K_S^0\eta$ . Dominant backgrounds stem from  $J/\psi \to \gamma K_S^0K_S^0\eta^0$  and  $J/\psi \to \gamma K_S^0K_S^0\eta^0\pi^0$ . These non- $\eta$  backgrounds are considered in the Partial Wave Analysis (PWA) by selecting events from data in the  $\eta$  sideband regions defined as  $0.45 < M_{\gamma\gamma} < 0.48~{\rm GeV}/c^2$  and  $0.60 < M_{\gamma\gamma} < 0.63~{\rm GeV}/c^2$ , and they account for about 2.5% of the total number of events in the  $\eta$  signal region.

A PWA of events satisfying  $M_{K_S^0K_S^0\eta} < 2.8~{\rm GeV}/c^2$  and  $M_{K_S^0K_S^0} < 1.1~{\rm GeV}/c^2$  is performed to determine the parameters of the structure around 1.85  ${\rm GeV}/c^2$ . These restrictions reduce complexities due to additional intermediate processes. The signal amplitudes are parameterized as sequential two-body decays, according to the isobar model:  $J/\psi \to \gamma X, \ X \to Y\eta$  or  $ZK_S^0$ , where Y and Z represent the  $K_S^0K_S^0$  and  $K_S^0\eta$  isobars, respectively. Parity conservation in the  $J/\psi \to \gamma K_S^0K_S^0\eta$  decay restricts the possible  $J^{PC}$  of the  $K_S^0K_S^0\eta$  (X) system to be  $0^{-+}$ ,  $1^{++}$ ,  $2^{++}$ ,  $2^{-+}$ ,  $3^{++}$ , etc.

In this letter, only spins J < 3 and possible S-wave or P-wave decays of the X are considered. The amplitudes are constructed using the covariant tensor formalism described in Ref. [16]. The relative magnitudes and phases of the partial wave amplitudes are determined by an unbinned maximum likelihood fit to data. The contribution of non- $\eta$  background events is accounted for in the fit by subtracting the negative log-likelihood (NLL) value obtained for events in the  $\eta$  sideband region from the NLL value obtained for events in the  $\eta$  signal region. The statistical significance of a contribution is estimated by the difference in NLL with and without the particular contribution, taking the change in degrees of freedom into account.

Our initial PWA fits include an X(1835) resonance in the  $f_0(980)\eta$  channel and a non-resonant component in one of the possible decay channels  $f_0(980)\eta$ ,  $f_0(1500)\eta$  or  $f_2(1525)\eta$ . All possible  $J^{PC}$  combinations of the X(1835)and the non-resonant component are tried. We then extend the fits by including an additional resonance at lower  $K_S^0 K_S^0 \eta$  mass. This additional component, denoted here as the X(1560), improves the fit quality when it is allowed to interfere with the X(1835). Our final fits show that the data can be best described with three components:  $X(1835) \rightarrow f_0(980)\eta$ ,  $X(1560) \rightarrow f_0(980)\eta$  and a non-resonant  $f_0(1500)\eta$  component. The  $J^{PC}$  of the X(1835), the X(1560) and the non-resonant component are all found to be  $0^{-+}$ . The X(1835), X(1560) and  $f_0(1500)$  are described by non-relativistic Breit-Wigner functions, where the intrinsic widths are not energy dependent. The masses and widths of the X(1835) and X(1560) are derived by scanning each over a certain range. The  $f_0(1500)$  mass and width are fixed to the values reported in Ref. [17]. The  $f_0(980)$  is parameterized by the Flatté formula [18], with the parameters fixed to the values reported by BESII [19] in the channels  $J/\psi \to \phi \pi^+ \pi^-$  and  $J/\psi \to \phi K^+ K^-$ . The scan returns a mass and width of the X(1835) of  $1844 \pm 9 \,\mathrm{MeV}/c^2$  and  $192^{+20}_{-17}$  MeV, respectively. The mass and width of the X(1560) are determined to be  $1565 \pm 8 \,\mathrm{MeV}/c^2$  and  $45^{+14}_{-13}$  MeV, respectively. Using a detection efficiency of 5.5%, obtained by a MC sample weighted by partial wave amplitudes, the product branching fraction of  $J/\psi \to \gamma X(1835)$  and  $X(1835) \to K_S^0 K_S^0 \eta$  ( $\mathcal{B}_{X(1835)}$ ) is calculated to be  $(3.31^{+0.33}_{-0.30}) \times 10^{-5}$ , where the decay  $X(1835) \to K_S^0 K_S^0 \eta$  is dominated by  $f_0(980)$  production. The  $K_S^0 K_S^0 \eta$ ,  $K_S^0 K_S^0$ ,  $K_S^0 \eta$  mass spectra and the distributions of the  $J/\psi$ ,  $K_S^0 K_S^0 \eta$  and  $K_S^0 K_S^0$  decay angles are shown in Fig. 2. Overlaid on the data are the PWA fit projections, as well as the individual contributions from each component. The  $\chi^2/n_{\rm bin}$  value is displayed on each figure to demonstrate the goodness of fit. We evaluate the significance by applying the likelihood ratio test, performing a separate fit for every systematic variation detailed below. The most conservative statistical significances of the X(1835) and X(1560) are  $12.9\sigma$  and  $8.9\sigma$ , respectively.

Various fits are performed by changing the  $J^{PC}$  and

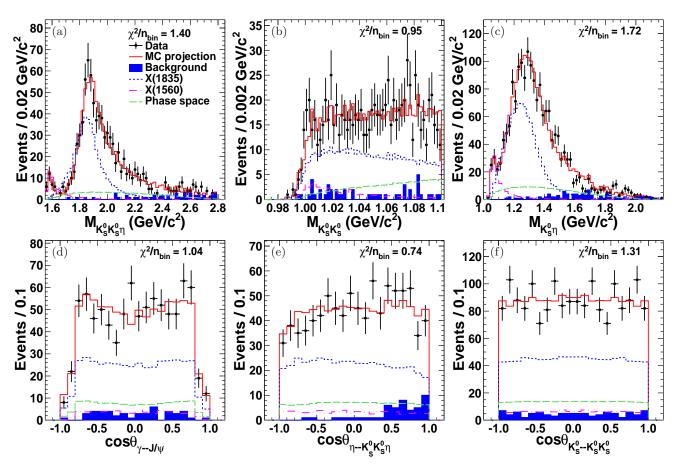


FIG. 2. Comparisons between data and PWA fit projections. (a), (b) and (c) are the invariant mass distributions of  $K_S^0 K_S^0 \eta$ ,  $K_S^0 K_S^0 \eta$  and  $K_S^0 \eta$  (two entries/event), respectively. (d)-(f) are the angular distributions of  $\cos \theta$ , where  $\theta$  is the polar angle of (d)  $\gamma$  in the  $J/\psi$  rest system; (e)  $\eta$  in the  $K_S^0 K_S^0 \eta$  rest system; and (f)  $K_S^0$  in the  $K_S^0 K_S^0 \eta$  rest system (two entries/event). The dots with error bars are data, the solid histograms are the PWA total projections, the shaded histograms are the non- $\eta$  backgrounds estimated by the  $\eta$  sideband, and the short-dashed, dash-dotted and long-dashed histograms show the contributions of X(1835), X(1560) and the non-resonant component, respectively.

decay mode of the non-resonant component compared to the nominal solution described above. The NLL value of a fit with a 1<sup>++</sup> non-resonant  $f_0(1500)\eta$  component is only worse by 0.8 compared to the nominal solution, which indicates that we cannot distinguish between the two spin assignments of the non-resonant component with our present statistics. This ambiguity introduces large systematic uncertainties in the  $\mathcal{B}_{X(1835)}$ , since the interference between the X(1835), X(1560) and the nonresonant component depends on the spin-assignment of the latter. To establish the  $J^{PC}$  of the X(1835), we perform a series of PWA fits assuming alternative  $J^{PC}$  hypotheses for both the X(1835) and the non-resonant contribution. For the non-resonant contribution, we also test several possible decay channels  $(f_0(980)\eta, f_0(1500)\eta)$  and  $f_2(1525)\eta$ ) in turn. For each non-resonant component assumption, the X(1835) 0<sup>-+</sup> hypothesis is significantly better than the  $1^{++}$  or  $2^{-+}$  hypotheses, with the NLL value improving by at least 41.6 units. Analogously, we perform the same series of PWA fits for the X(1560). Again the  $0^{-+}$  hypothesis for the X(1560) always yields

a significantly better fit result than other  $J^{PC}$  assignments, with the NLL value improving by at least 12.8 units.

We evaluate the contributions from additional wellknown resonances by adding them individually to the We consider all possible combinations for Xand its subsequent decay products Y and Z as given in Ref. [17]: for X, this includes  $\eta(1760)$ ,  $\eta(2225)$ ,  $f_1(1510), \quad \eta_2(1870), \quad f_2(1810), \quad f_2(1910), \quad f_2(1950),$  $f_2(2010), f_2(2150), f_2(2300), f_2(2340), f_J(2220);$  for Y,  $f_0(980), f_0(1500), f_0(1710), f_2(1270)$  and  $f_2(1525)$ ; for  $Z, K^*(1410), K^*(1680), K_0^*(1430), K_0^*(1950), K_2^*(1430)$ and  $K_2^*$  (1980). Additional non-resonant contributions with various  $J^{PC}$  and decay modes are studied as well. The statistical significances of the additional contributions are smaller than  $5\sigma$ . In order to check the possible contribution from a non-resonant  $K_S^0 K_S^0$  process, we add a  $X(1835) \to (K_S^0 K_S^0)_S \eta$  process into the nominal solution, where  $(K_S^0 K_S^0)_S$  refers to a non-resonant  $K_S^0 K_S^0$  contribution in a relative S-wave. We find the resulting significance of the  $X(1835) \rightarrow f_0(980)\eta$  process and the  $X(1835) \to (K_S^0 K_S^0)_S \eta$  process to be  $6.8\sigma$  and  $1.6\sigma$ , respectively, so we do not include the latter process in the nominal solution. We also test a fit by changing the decay mode of the X(1560) in the nominal solution from  $f_0(980)\eta$  to  $(K_S^0 K_S^0)_S \eta$ ; the fit with  $X(1560) \to (K_S^0 K_S^0)_S \eta$  has almost the same fit quality as the nominal solution. Therefore, with the present statistics, we cannot draw a conclusion about the X(1560) decay mode. The largest differences in masses and widths of the X(1835) and X(1560) and the product branching fraction  $\mathcal{B}_{X(1835)}$  between all above alternative fits and the nominal solution are taken as systematic uncertainties from the components in the nominal solution.

For the measurements of the masses and widths of the X(1835) and X(1560) and the product branching fraction  $\mathcal{B}_{X(1835)}$ , we include the following sources of systematic uncertainties in addition to the sources discussed above: we change the  $K^0_S K^0_S$  mass range to  $M_{K^0_S K^0_S} <$ 1.05, 1.15 and 1.20 GeV/ $c^2$ ; we change the  $f_0(980)$ mass and coupling constants in the Flatté formula to other experimental measurements [20–22]; we change the  $f_0(1500)$  mass and width by one standard deviation [17]; we increase and decrease the non- $\eta$  background level by one standard deviation; we change the parameterization of the X(1835) and X(1560) line shape to a Breit-Wigner function whose intrinsic width is energy-dependent [23]; and we replace the X(1560) by  $\eta(1405)$  or  $\eta(1475)$ . For the systematic errors of the product branching fraction  $\mathcal{B}_{X(1835)}$ , we also consider the following additional uncertainties. The  $K^0_S$  reconstruction efficiency is studied using two control samples of  $J/\psi \to K^{*\pm}K^{\mp}$  and  $J/\psi \to \phi K_S^0 K^{\pm} \pi^{\mp}$ , while the photon detection efficiency is investigated based on a clean sample of  $J/\psi \to \rho \pi$ . The differences between data and MC simulation are 1.0% for each  $K_S^0$  and 1.0% for each photon [24]. A control sample of  $J/\psi \to \gamma K_S^0 K_S^0 \pi^0$  is selected to estimate the uncertainty associated with the 4C kinematic fit. The efficiency is the ratio of the signal yields with and without the kinematic fit requirement  $\chi^2_{4C}$  < 40. The difference between data and MC, 1.5%, is assigned as the systematic uncertainty. We also consider the uncertainties from the number of  $J/\psi$  events [10, 11] and the branching fractions of  $K_S^0 \to \pi^+\pi^-$  and  $\eta \to \gamma\gamma$  [17]. We change the mass and width of X(1835) or X(1560) by one standard deviation of the statistical uncertainty. The individual uncertainties are assumed to be independent and are added in quadrature to obtain the total systematic uncertainties as presented in the Supplemental Material [25].

In summary, a PWA of  $J/\psi \to \gamma K_S^0 K_S^0 \eta$  has been performed in the mass range  $M_{K_S^0 K_S^0 \eta} < 2.8 \ {\rm GeV}/c^2$  after requiring  $M_{K_S^0 K_S^0} < 1.1 \ {\rm GeV}/c^2$ . The PWA fit requires a contribution from  $X(1835) \to K_S^0 K_S^0 \eta$  with a statistical significance greater than  $12.9\sigma$ , where the

 $X(1835) \to K_S^0 K_S^0 \eta$  is dominated by  $f_0(980)$  production. The spin-parity of the X(1835) is determined to be  $0^{-+}$ . The mass and width of the X(1835) are measured to be  $1844 \pm 9(\text{stat})_{-25}^{+16}(\text{syst}) \,\text{MeV}/c^2$  and  $192_{-17}^{+20}(\text{stat})_{-43}^{+62}(\text{syst}) \,\text{MeV}$ , respectively. The corresponding product branching fraction  $\mathcal{B}_{X(1835)}$  is measured to be  $(3.31_{-0.30}^{+0.33}(\text{stat})_{-1.29}^{+1.96}(\text{syst})) \times 10^{-5}$ . The mass and width of the X(1835) are consistent with the values obtained from the decay  $J/\psi \to \gamma \pi^+ \pi^- \eta'$  by BESIII [2]. These results are all first-time measurements and provide important information to further understand the nature of the X(1835).

Another  $0^{-+}$  state, the X(1560), also is observed in data with a statistical significance larger than  $8.9\sigma$  and is seen to interfere with the X(1835). The mass and width of the X(1560) are determined to be  $1565\pm 8(\mathrm{stat})^{+0}_{-63}(\mathrm{syst})\,\mathrm{MeV}/c^2$  and  $45^{+14}_{-13}(\mathrm{stat})^{+21}_{-28}(\mathrm{syst})\,\mathrm{MeV}$ , respectively. The mass and width of the X(1560) are consistent with those of the  $\eta(1405)$  and  $\eta(1475)$  as given in Ref. [17] within  $2.0\sigma$  and  $1.4\sigma$ , respectively. Present statistics do not allow us to conclusively determine if the X(1560) is the same state as the  $\eta(1405)/\eta(1475)$  or a new meson. More statistics in this analysis and an amplitude analysis of  $J/\psi \to \gamma \eta \pi^0 \pi^0$  and  $J/\psi \to \gamma K_S^0 K_S^0 \pi^0$  processes may help to understand the nature of the X(1560).

The BESIII collaboration thanks the staff of BEPCII and the IHEP computing center for their strong sup-This work is supported in part by National Key Basic Research Program of China under Contract No. 2015CB856700; National Natural Science Foundation of China (NSFC) under Contracts Nos. 11125525, 11235011, 11322544, 11335008, 11425524; the Chinese Academy of Sciences (CAS) Large-Scale Scientific Facility Program; the CAS Center for Excellence in Particle Physics (CCEPP); the Collaborative Innovation Center for Particles and Interactions (CICPI); Joint Large-Scale Scientific Facility Funds of the NSFC and CAS under Contracts Nos. 11179007, U1232201, U1332201; CAS under Contracts Nos. KJCX2-YW-N29, KJCX2-YW-N45; 100 Talents Program of CAS; INPAC and Shanghai Key Laboratory for Particle Physics and Cosmology; German Research Foundation DFG under Contract No. Collaborative Research Center CRC-1044; Istituto Nazionale di Fisica Nucleare, Italy; Ministry of Development of Turkey under Contract No. DPT2006K-120470; Russian Foundation for Basic Research under Contract No. 14-07-91152; U.S. Department of Energy under Contracts Nos. DE-FG02-04ER41291, DE-FG02-05ER41374, DE-FG02-94ER40823, DESC0010118; U.S. National Science Foundation; University of Groningen (RuG) and the Helmholtzzentrum fuer Schwerionenforschung GmbH (GSI), Darmstadt; WCU Program of National Research Foundation of Korea under Contract No. R32-2008-000-10155-0.

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